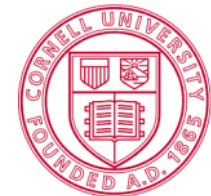


Microscale Atmospheric Re-Entry Sensors

Justin A. Atchison
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Cornell University
Space Systems Design Studio



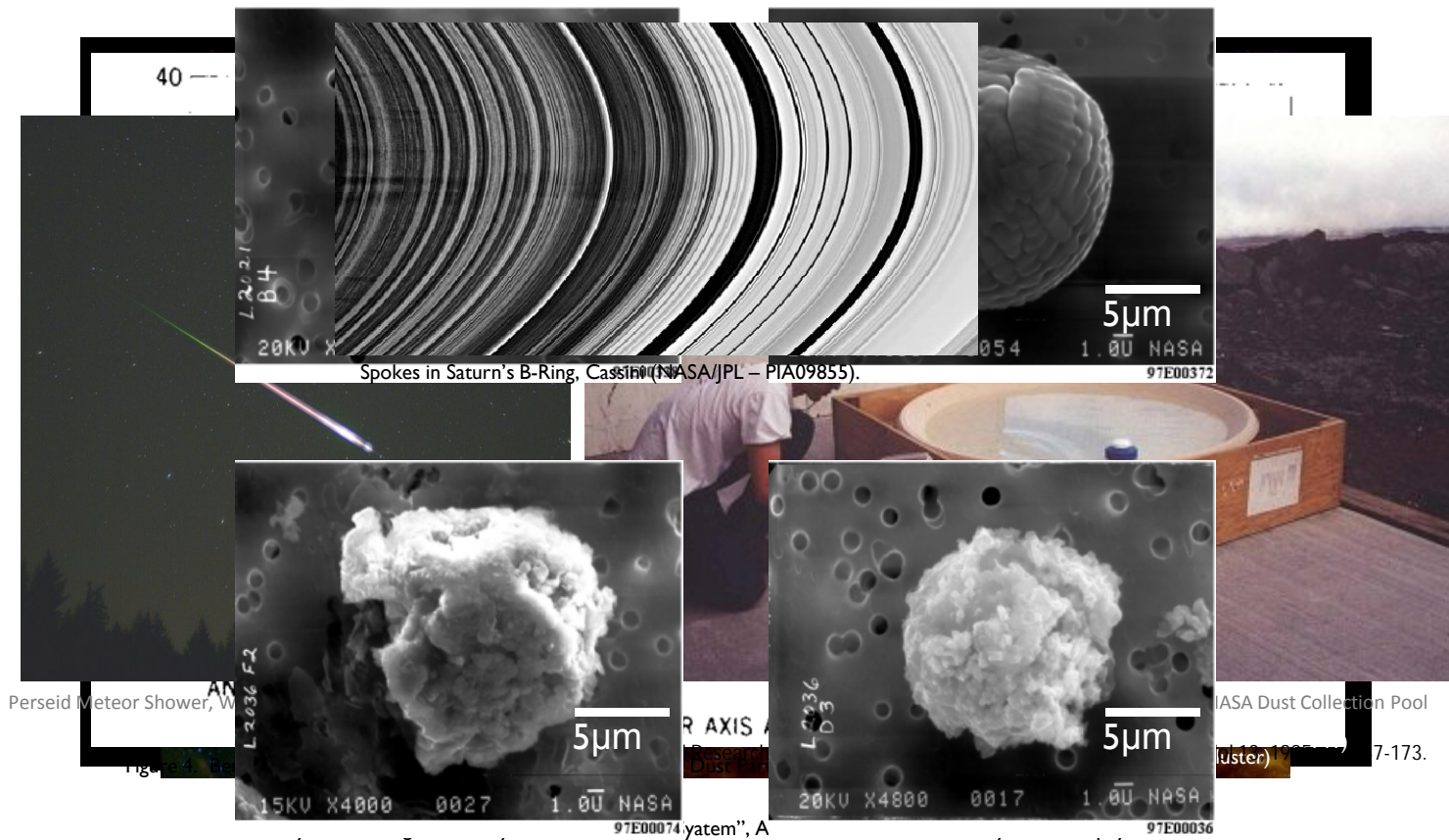
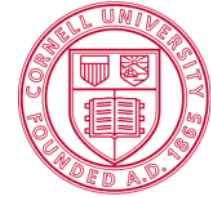
Co-Authors: Zachary R. Manchester
Dr. Mason A. Peck

International Planetary Probe Workshop - 7
June 14-18 2010

Dust Dynamics

Cornell University

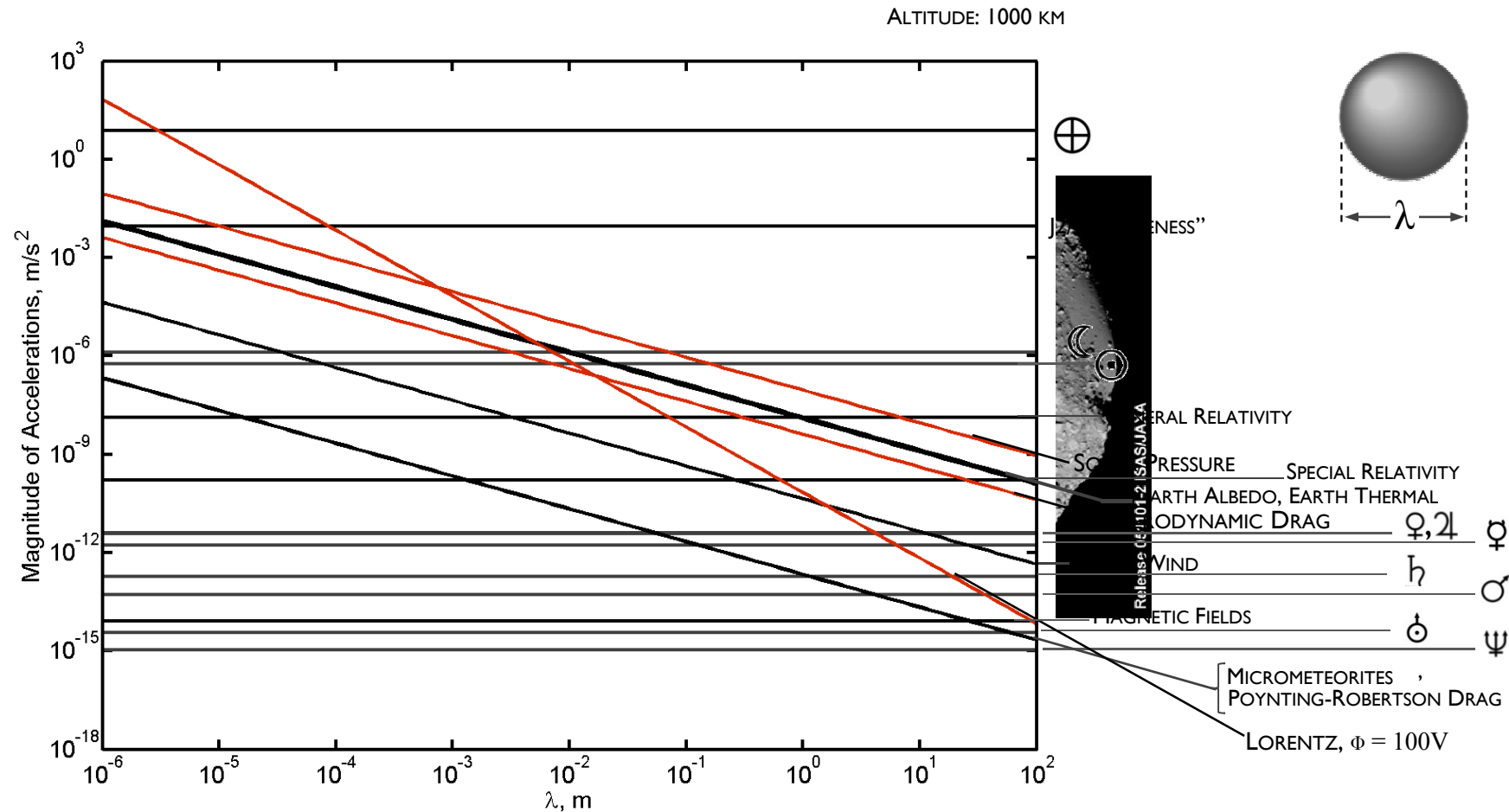
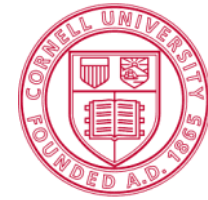
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Many Accelerations Depend on Length-Scale

Cornell University

Space Systems Design Studio



Images Courtesy NASA / JAXA

Concept

Design

Simulation

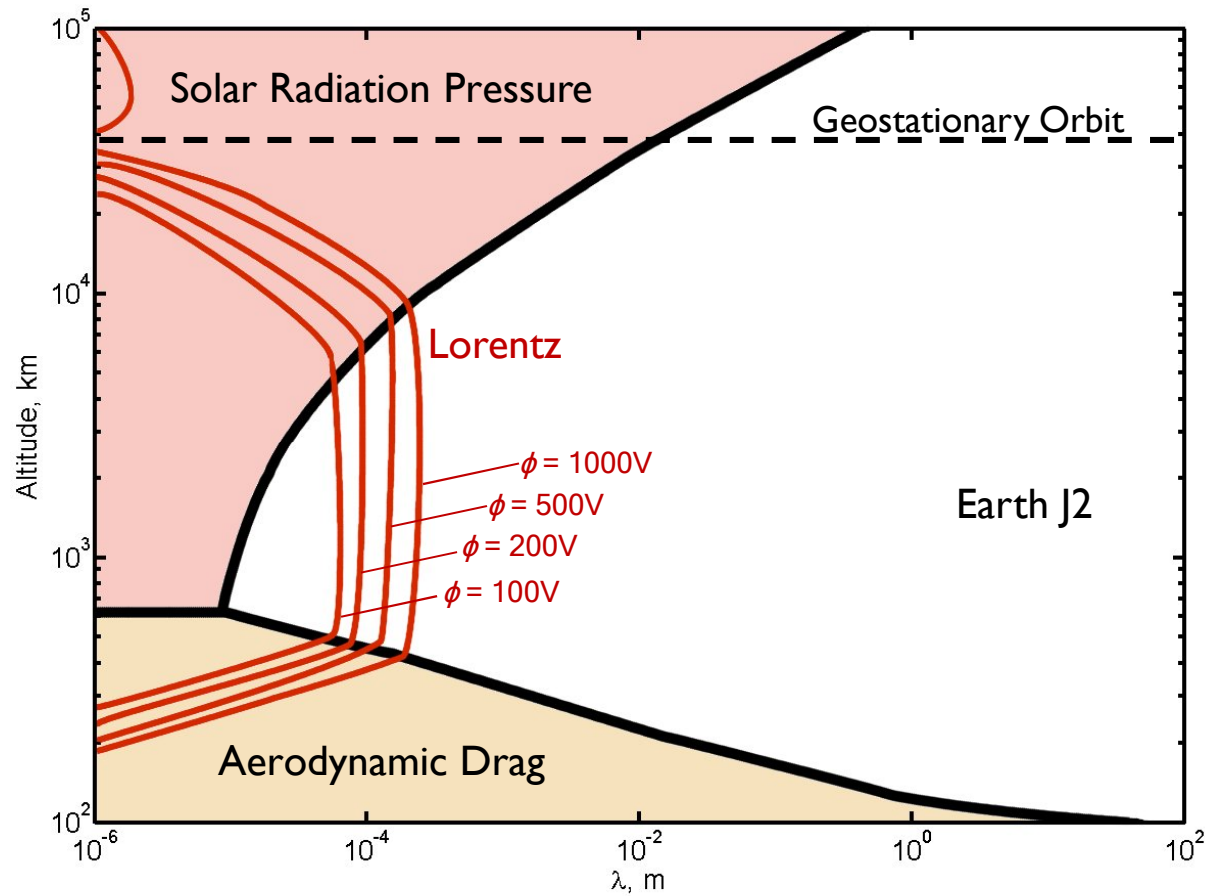
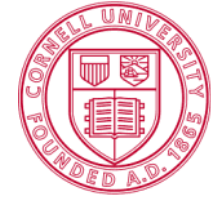
Conclusions

JAA - 3 / 19

Length-Scale and Altitude

Cornell University

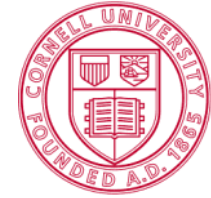
Space Systems Design Studio



Area-to-Mass and Characteristic Length

Cornell University

Space Systems Design Studio



$$a = \frac{F}{m} \propto \frac{A}{m} \quad \begin{array}{l} \swarrow A = \kappa_A \lambda^2 \\ \nwarrow m = \kappa_V \rho \lambda^3 \end{array} \quad \frac{A}{m} = \frac{1}{\rho} \frac{\kappa_A}{\kappa_V} \frac{1}{\lambda}$$

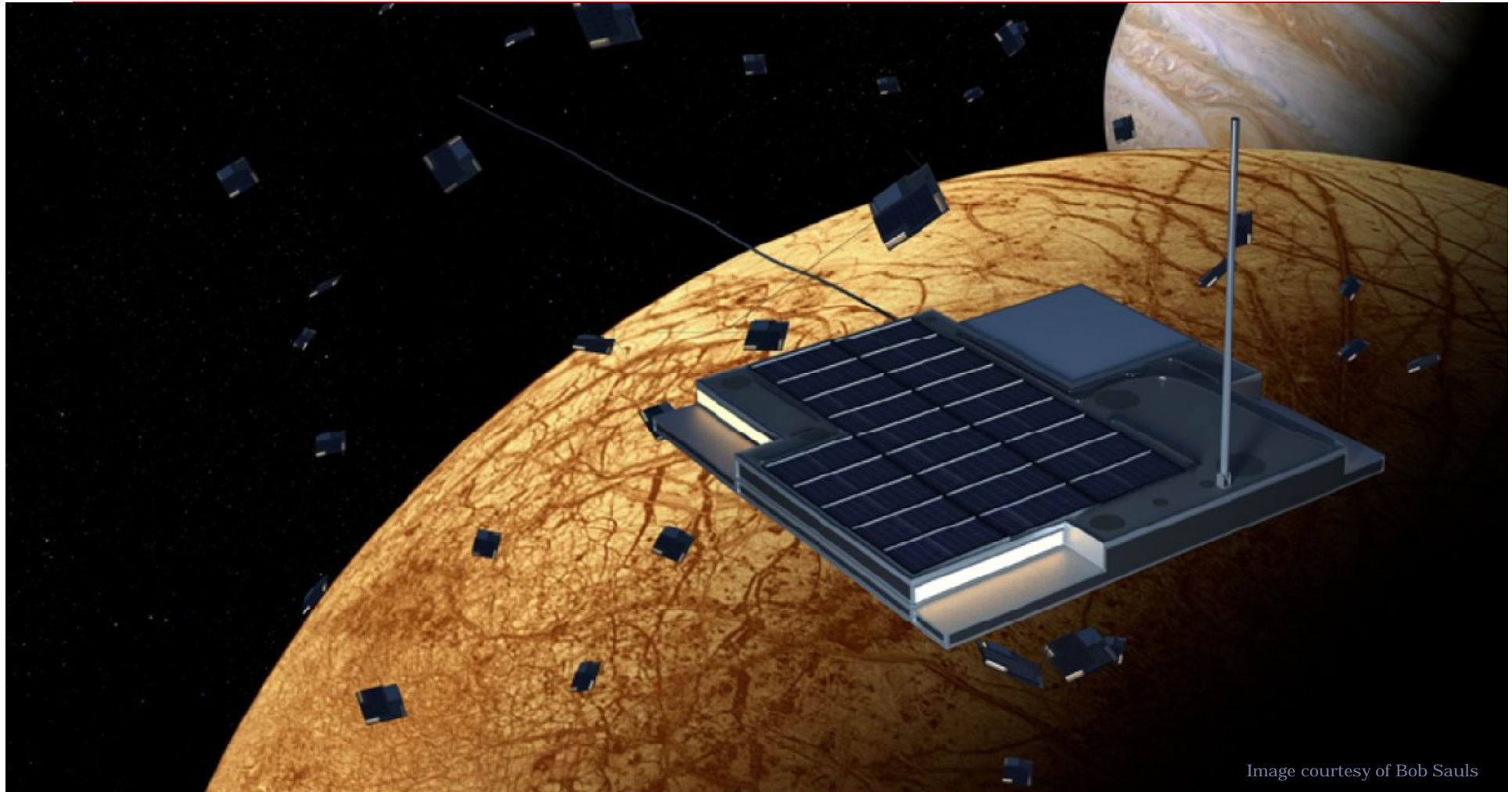
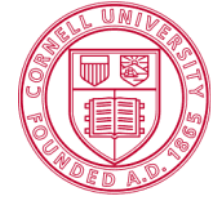
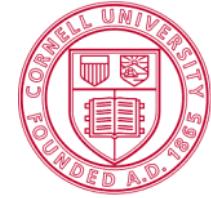


Image courtesy of Bob Sauls

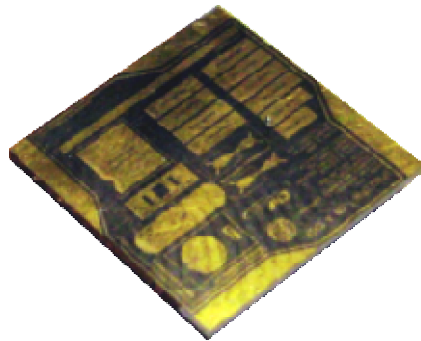
Spacecraft-on-Chip Architecture

Cornell University

Space Systems Design Studio



- Application specific integrated circuit (ASIC) with spacecraft functionality: power, communications, data handling, etc.
- Extremely thin form factor



$$\lambda = 1 \text{ cm}$$

$$\varepsilon = 0.0025$$

$$\rho = 2300 \text{ kg/m}^3, \text{ Silicon}$$

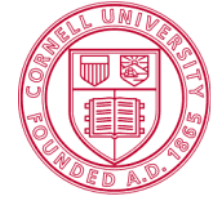
$$A = 1 \text{ cm}^2$$

$$m = 8 \text{ mg}$$

$$A/m = 125 \text{ m}^2/\text{kg}$$

$$\beta = 0.036$$

Sprite – A small, elusive supernatural being.



Traditional Motivation

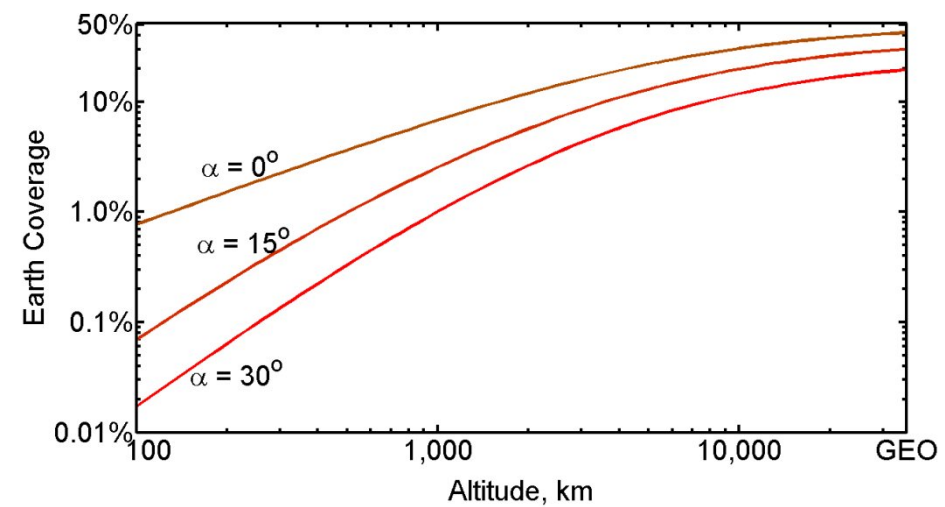
$$\text{Cost} = f(\text{Size, Mass, Power})$$

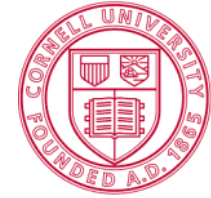
New Capabilities

"Can't we make 'Palm Pilot' versions of spacecraft instead of main frames?"

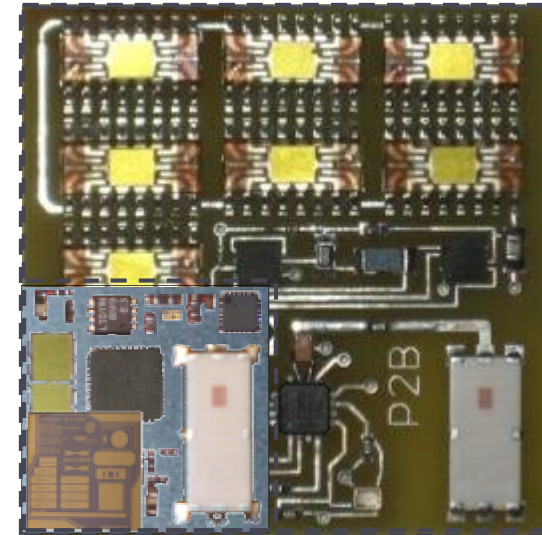
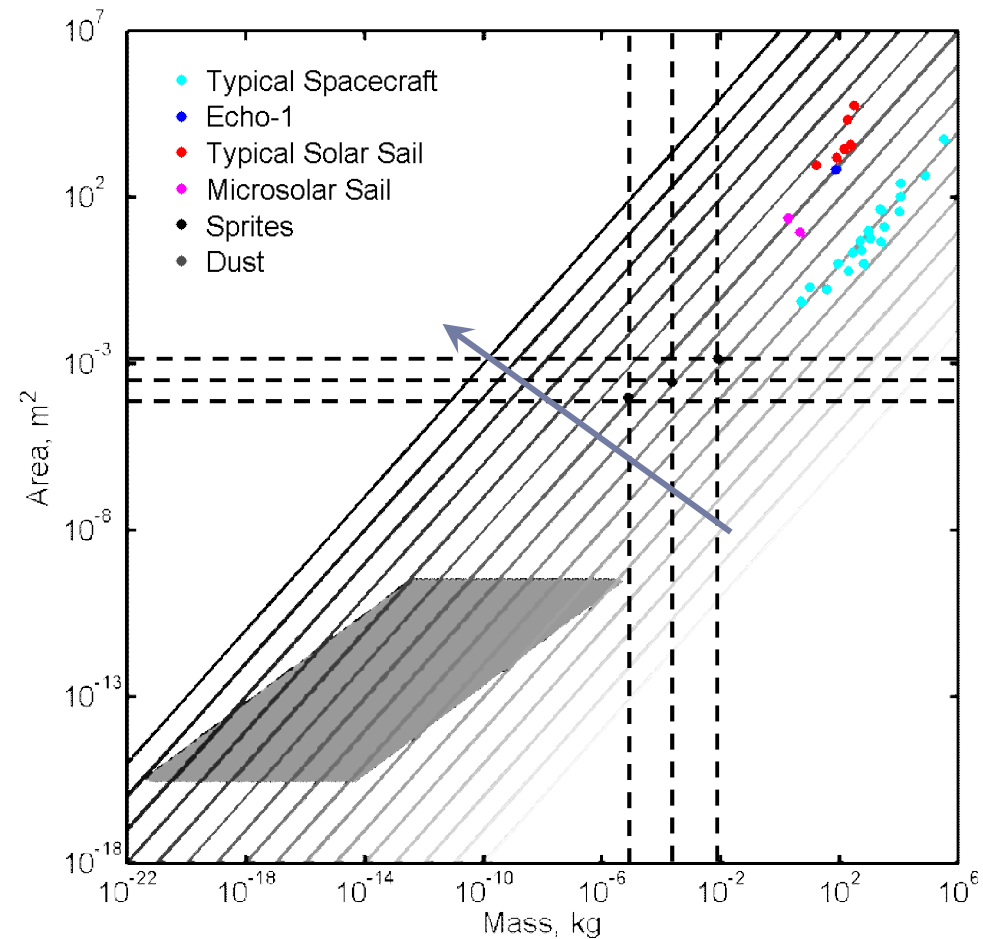
Siegfried W. Janson
Aerospace Corporation, May 2000

Distributed Sensing



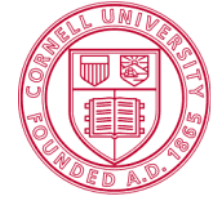


Area-to-Mass Survey

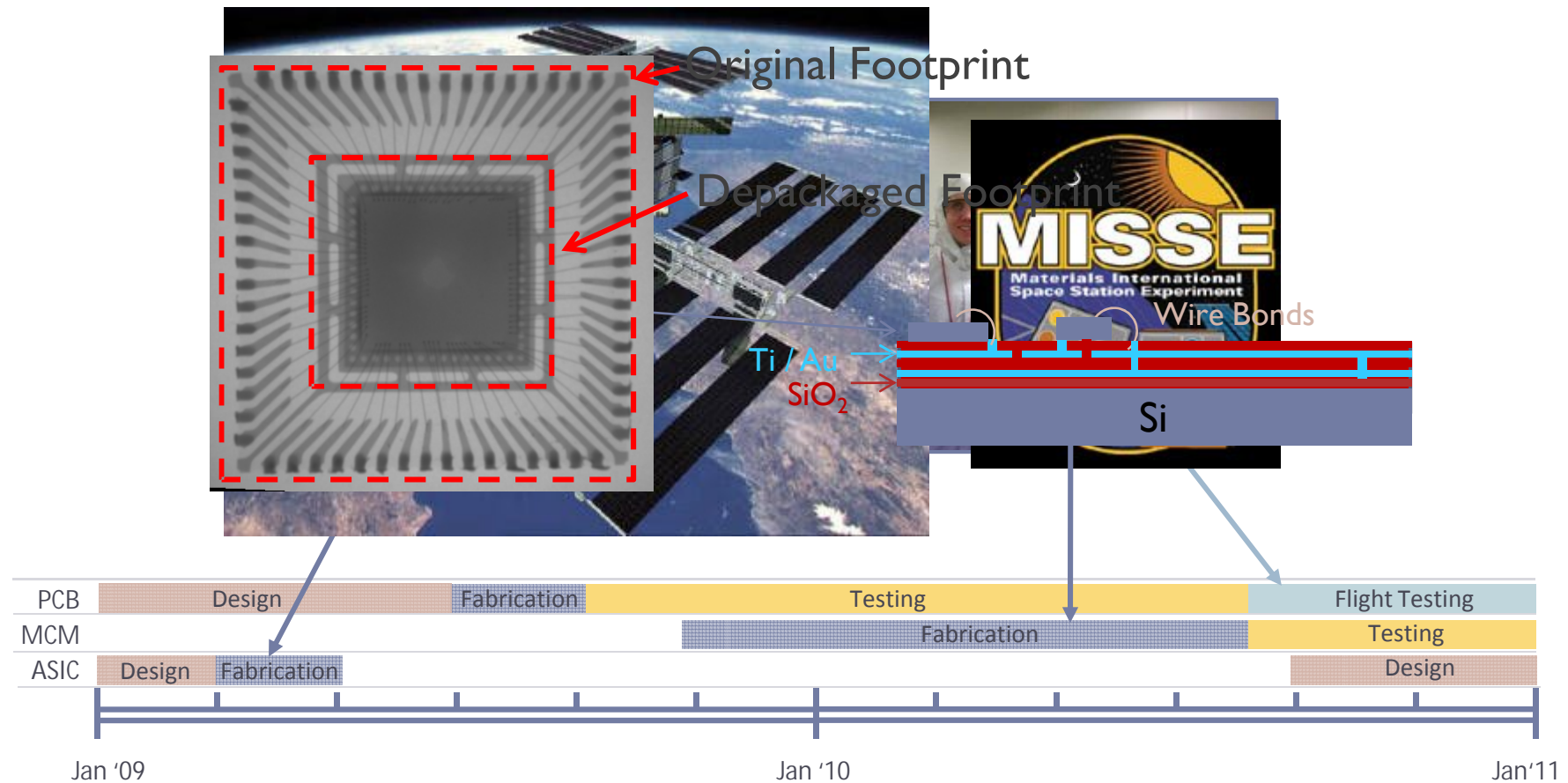


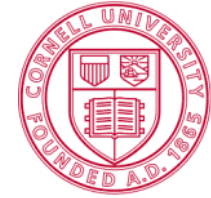
Sprite – Applied Circuits Board IC
 $A = 0.00049 \text{ m}^2$
 $m = 0.0000028 \text{ kg}$
 $A/m = 0.33 \text{ m}^2/\text{kg}$

*All Images Courtesy NASA

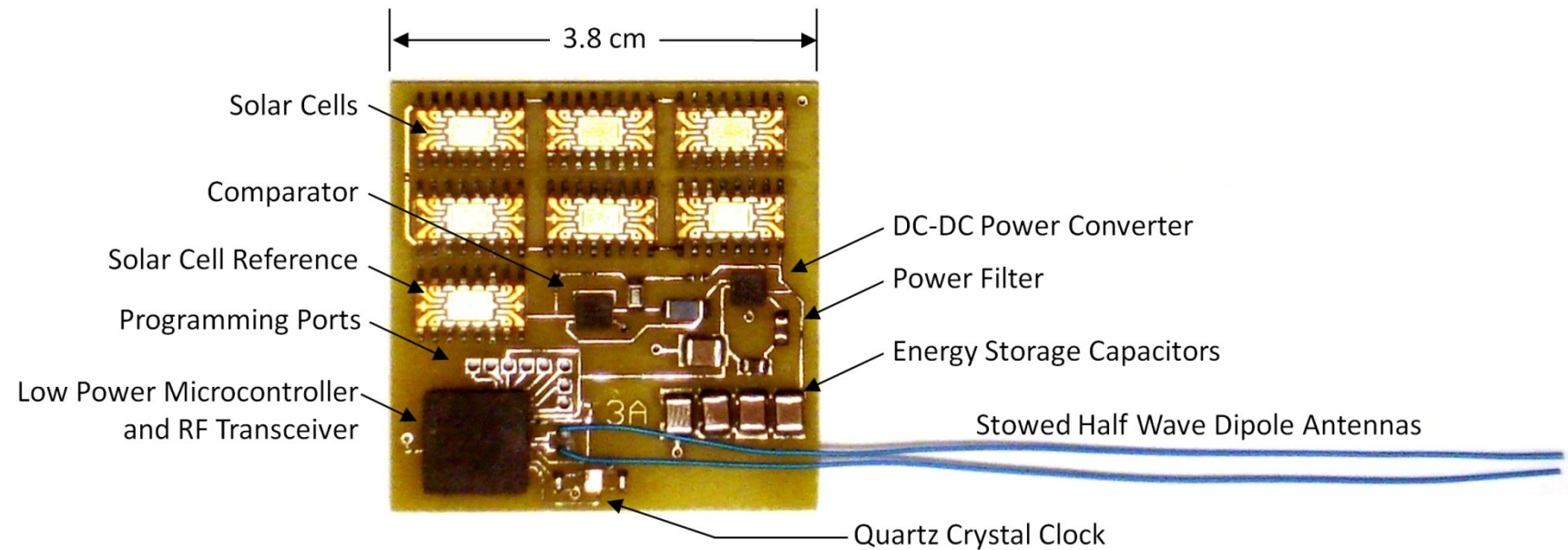


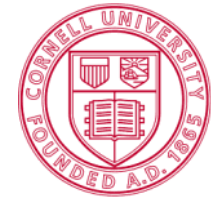
Sprite Prototypes





Sprite Features

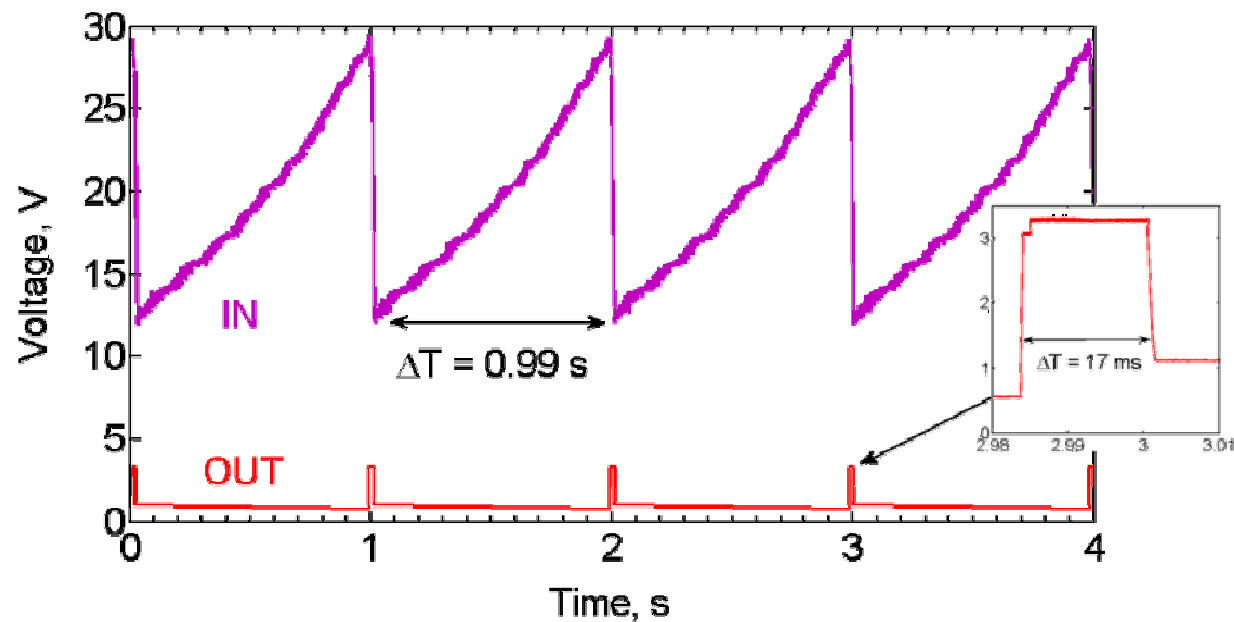


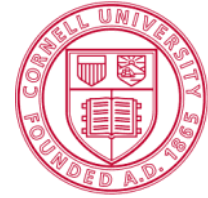


Power Subsystem

“Bursty” Operation

Solar cells charge a capacitor bank until there is sufficient energy to operate the microcontroller and radio.

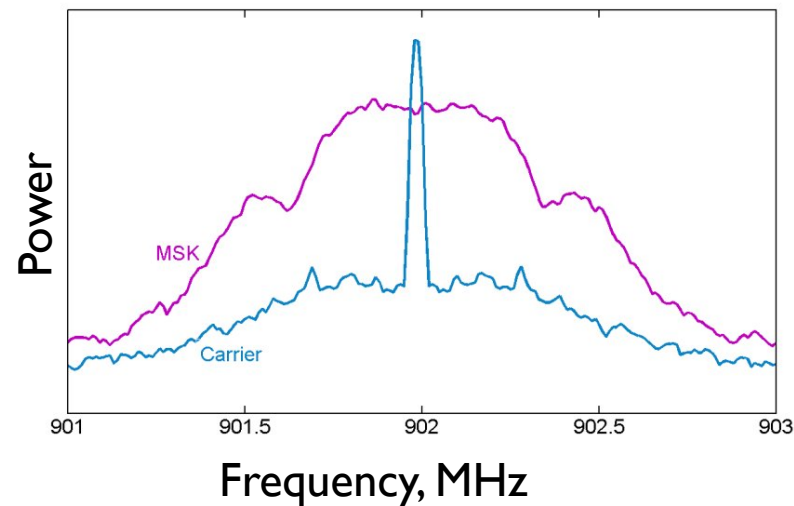




Communications

Matched Filtering

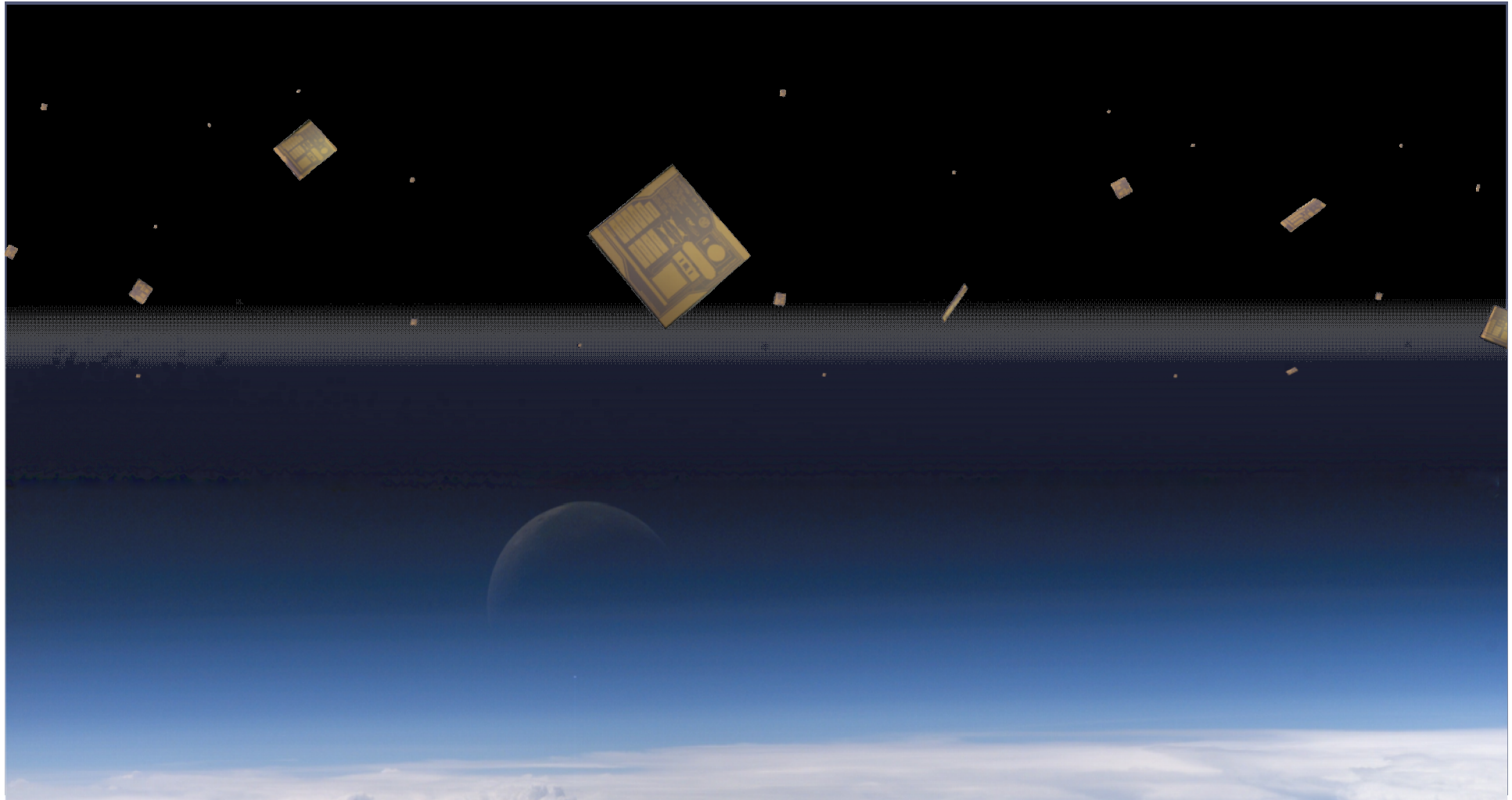
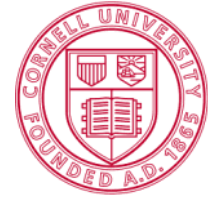
- The radio supplies only 12 mW
- 512 bit pseudorandom noise (PRN) sequence key for matched filtering offers ~30 dB of effective gain
- Minimum-Shift-Keying (MSK) for lower bit error rates



Atmospheric Sensors?

Cornell University

Space Systems Design Studio



Concept

Design

Simulation

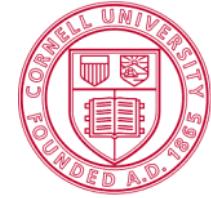
Conclusions

JAA - 14 / 19

How Does Dust Survive Re-Entry?

Cornell University

Space Systems Design Studio



Mechanical Forces*

- Magnitude of drag depends explicitly on A/m .
- Terminal velocity depends explicitly on A/m .
- Ram pressure goes with $\rho_A v^2$.

High-Temperature Ablation**

- **There is sufficient kinetic energy to ablate.**
- **Small grains achieve uniform temperature quickly.**
- Aerothermal heat-input goes with $\rho_A v^3$.
- Aerothermal heat-input is a function of the local flow properties.

Dust survives re-entry by shedding the majority of its kinetic energy in the rarefied atmosphere.

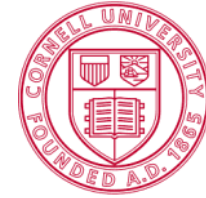
*Brownlee, D.E., Ann. Rev. Earth Planet. Sci. 1985.

**Chapman and Ferris, J. Geophys. Res., 1961.

Preliminary Aerothermal Model

Cornell University

Space Systems Design Studio



A Skin Friction Coefficient C_F encapsulates the flow regime

- Free-Molecular Flow*** ($10 \leq Kn$)
- Transitional *** ($0.01 \leq Kn < 10$)
- Continuum** ($Kn \leq 0.01$)
 - Laminar ($Re \leq 2300$)
 - Turbulent ($2300 < Re$)

$$Kn = \frac{\zeta_A}{\lambda} \quad Re = \frac{\rho_A v \lambda}{\mu_A}$$

$$\frac{dT}{dt} = \left(\frac{A}{m c_s} \right) \left[\underbrace{\frac{1}{4} \rho_{atm} C_F (Kn, Re) v^3}_{\text{Hypersonic Aerothermal Convection*}} - \underbrace{2\xi\sigma(T^4 - T_s^4)}_{\text{Radiation to Environment}} \right]$$

*Allen, H.J., and Eggers, A.J., "A Study of the Motion and Aerodynamic Heating of Missiles Entering the Earth's Atmosphere at High Supersonic Speeds," NASA Technical Report #1381, 1958.

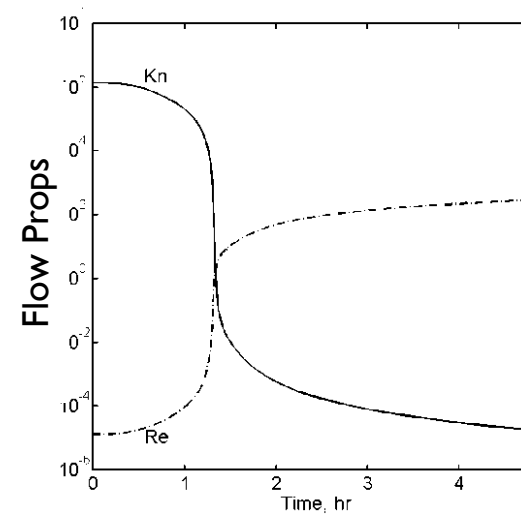
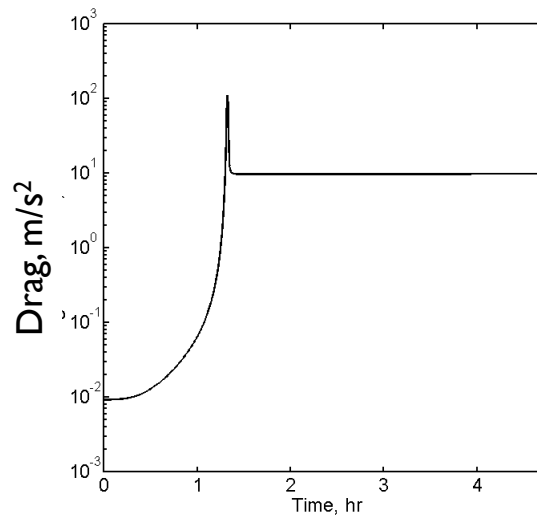
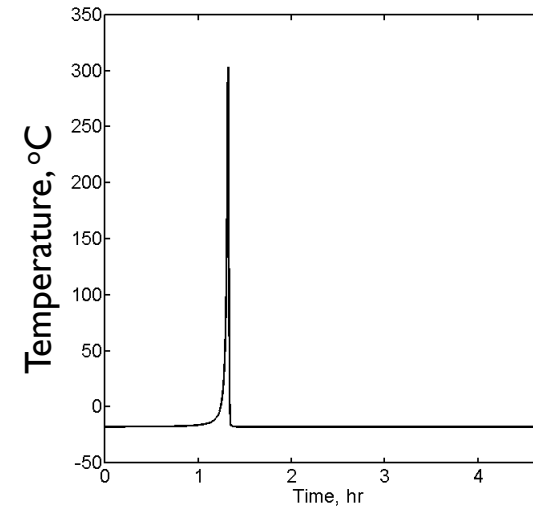
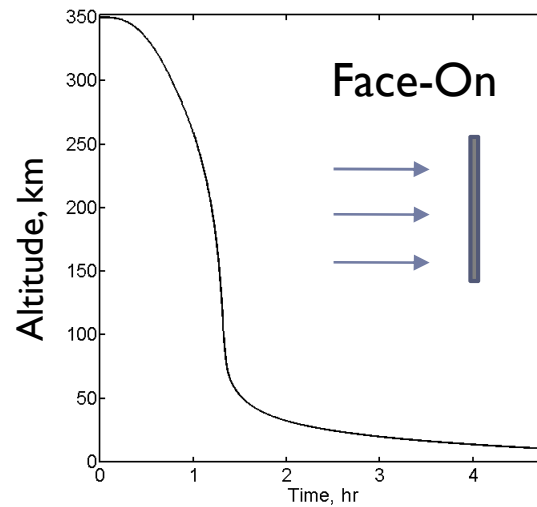
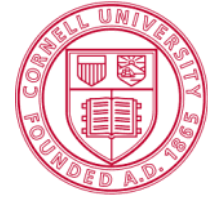
**Hirshcel, E.H., Basics of Aerothermodynamics, 1st Ed, Springer-Verlag, Berlin, 2005.

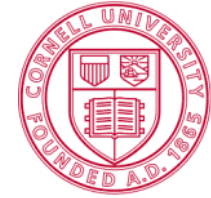
***Koppenwallner, G., Fritsche, Band Bolts of Disintegrating Spacecraft during Uncontrolled re-entry", Proceedings of the Third European Conference on Space Debris, 19 - 21 March 2001, .. and T. Lips, "Survivability and Ground Risk Potential of Screws Darmstadt, Germany.

Ballistic Re-Entry Simulations

Cornell University

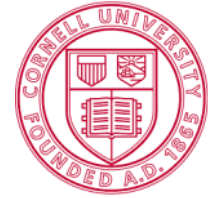
Space Systems Design Studio





Conclusions

- Area-to-Mass is a design parameter that can drive spacecraft dynamics.
- For the right choice of area-to-mass and characteristic length, a body can re-enter while maintaining a low temperature.
- A spacecraft-on-chip architecture may be capable of demonstrating this low-temperature behavior.
- I welcome your advice on how to move forward with this analysis.



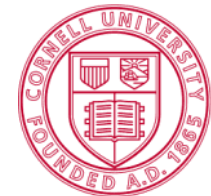
Acknowledgements

- Zac Manchester
- Matt Blair
- Victoria Alexander
- Robert MacCurdy
- Bernardo Cordovez
- Ryan Zhou
- Phillipe Tosi
- Sharon Kotz
- Alfred Ernst
- Space Systems Design Studio
- Dr. David Caughey
- Dr. Joseph Burns
- Dr. Brian Gilchrist of the University of Michigan
- Dr. Ben Harris of the University of Texas
- US National Science Foundation
- US Department of Defense

Microscale Atmospheric Re-Entry Sensors

Justin A. Atchison
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Space Systems Design Studio



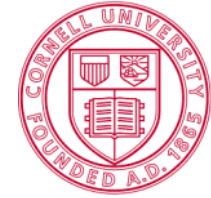
Co-Authors: Zachary R. Manchester
Dr. Mason A. Peck

International Planetary Probe Workshop - 7
June 14-18 2010

Ballistic Re-Entry Simulations

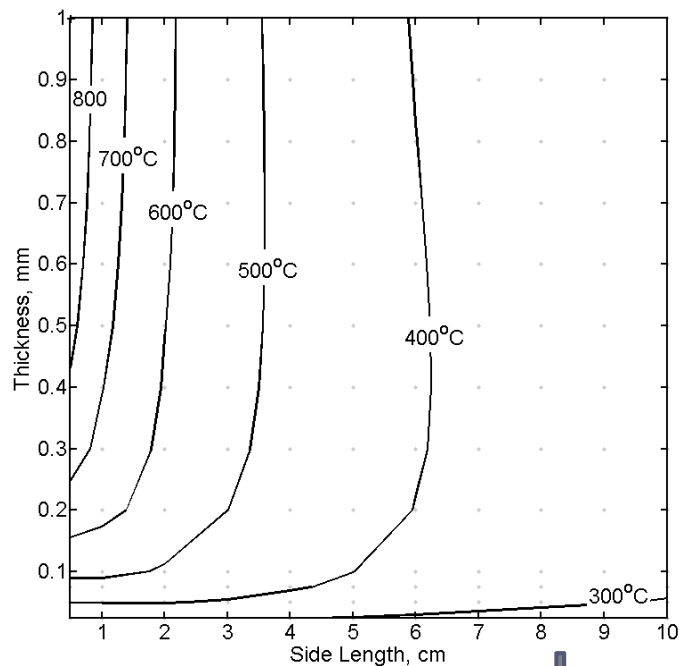
Cornell University

Space Systems Design Studio

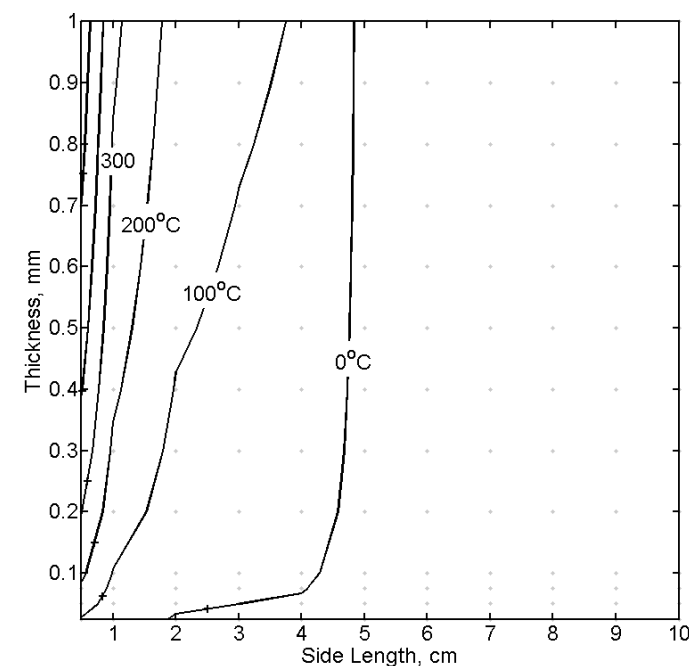
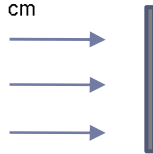


Parametric Search

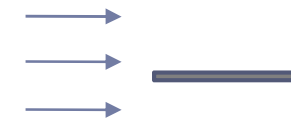
Length and Thickness vs. Maximum Temperature



Face-On



Edge-On



Concept

Design

Simulation

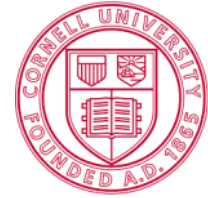
Conclusions

↳ Backup Slides

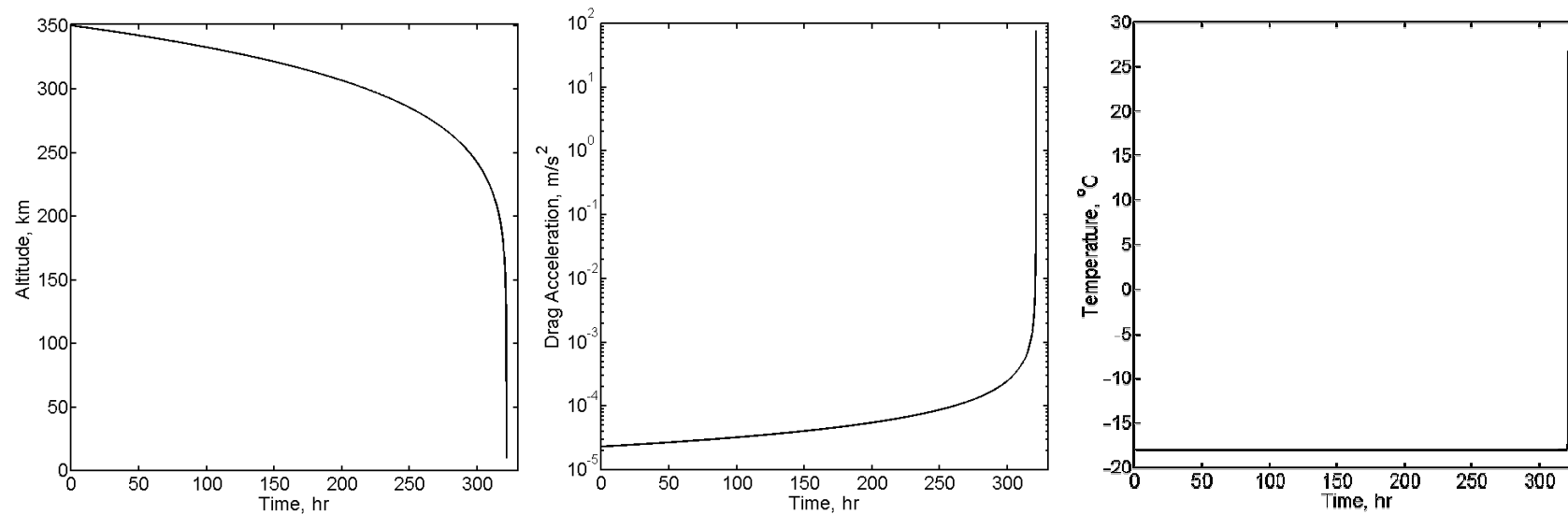
Ballistic Re-Entry Simulations

Cornell University

Space Systems Design Studio



Sample Re-Entry Simulation Edge-On, 1cm x 25 μ m



Concept

Design

Simulation

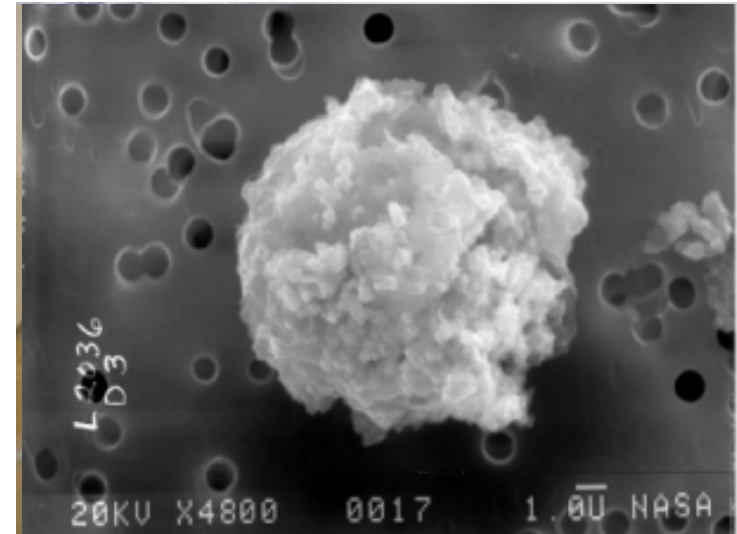
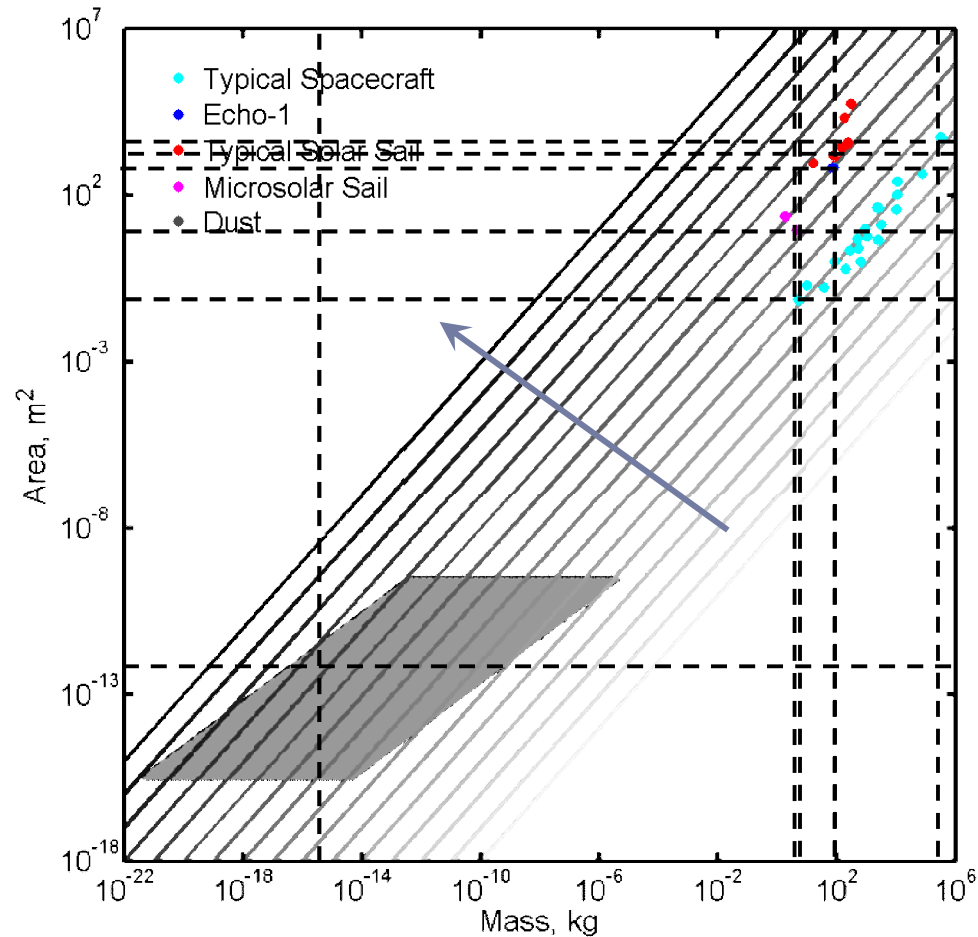
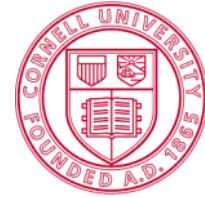
Conclusions

↳ Backup Slides

Survey of Area-to-Mass Values

Cornell University

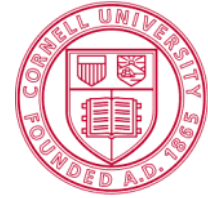
Space Systems Design Studio



~~RESTRICTED~~ Ball's Bluff Station

$$A = 0.00000000000078 \text{ m}^2$$
[illegible]
$$A/m = 0.00002712 \text{ kg}$$

*All Images Courtesy NASA



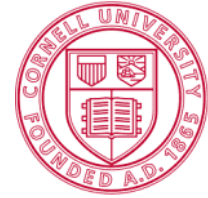
Aerodynamic Drag

- Drag
 - Often characterized by the ballistic coefficient.
 - Typical spacecraft values fall between 10 and 100 kg/m².
 - A 1cm x 1cm x 25 μm silicon wafer achieves $\beta = 0.02$.
 - Ballistic coefficient drives terminal velocity.

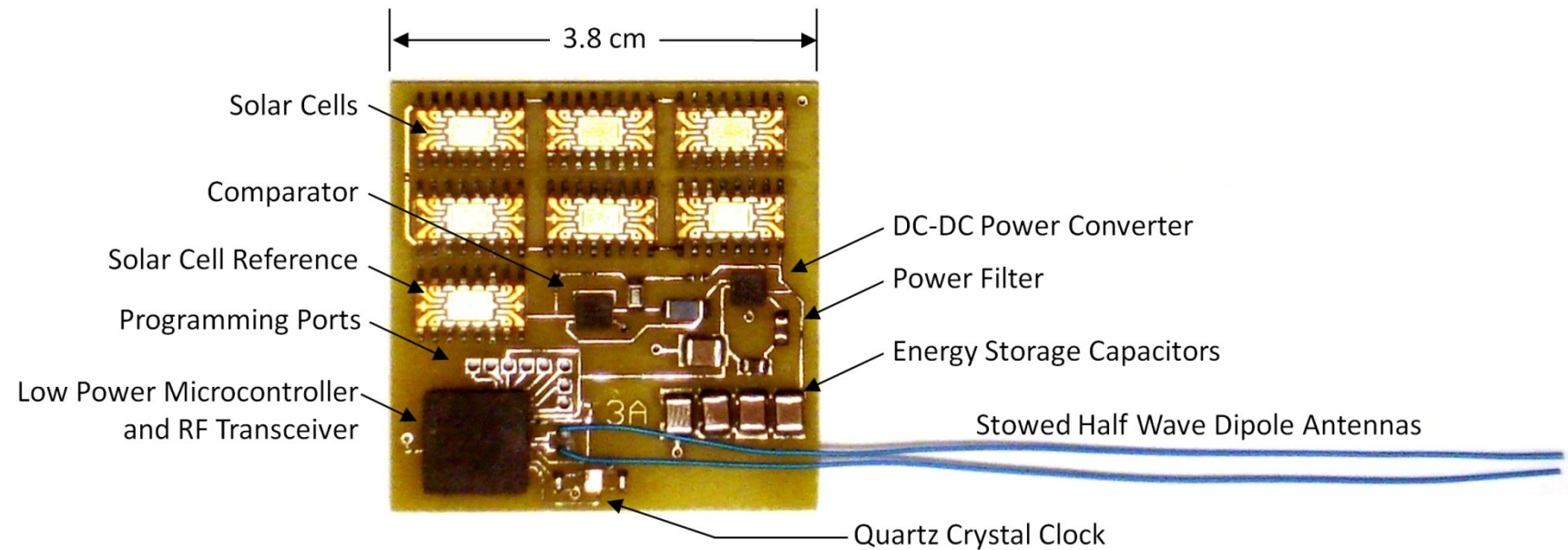
$$\mathbf{a}_{AD} = -\frac{1}{2} \frac{A}{m} C_D \rho_A v^2 \hat{\mathbf{v}}$$

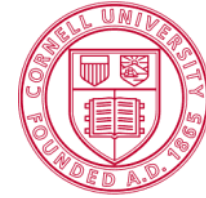
$$\beta = \frac{m}{AC_d}$$

$$v_T = \sqrt{\beta \frac{2g}{\rho_A}}$$



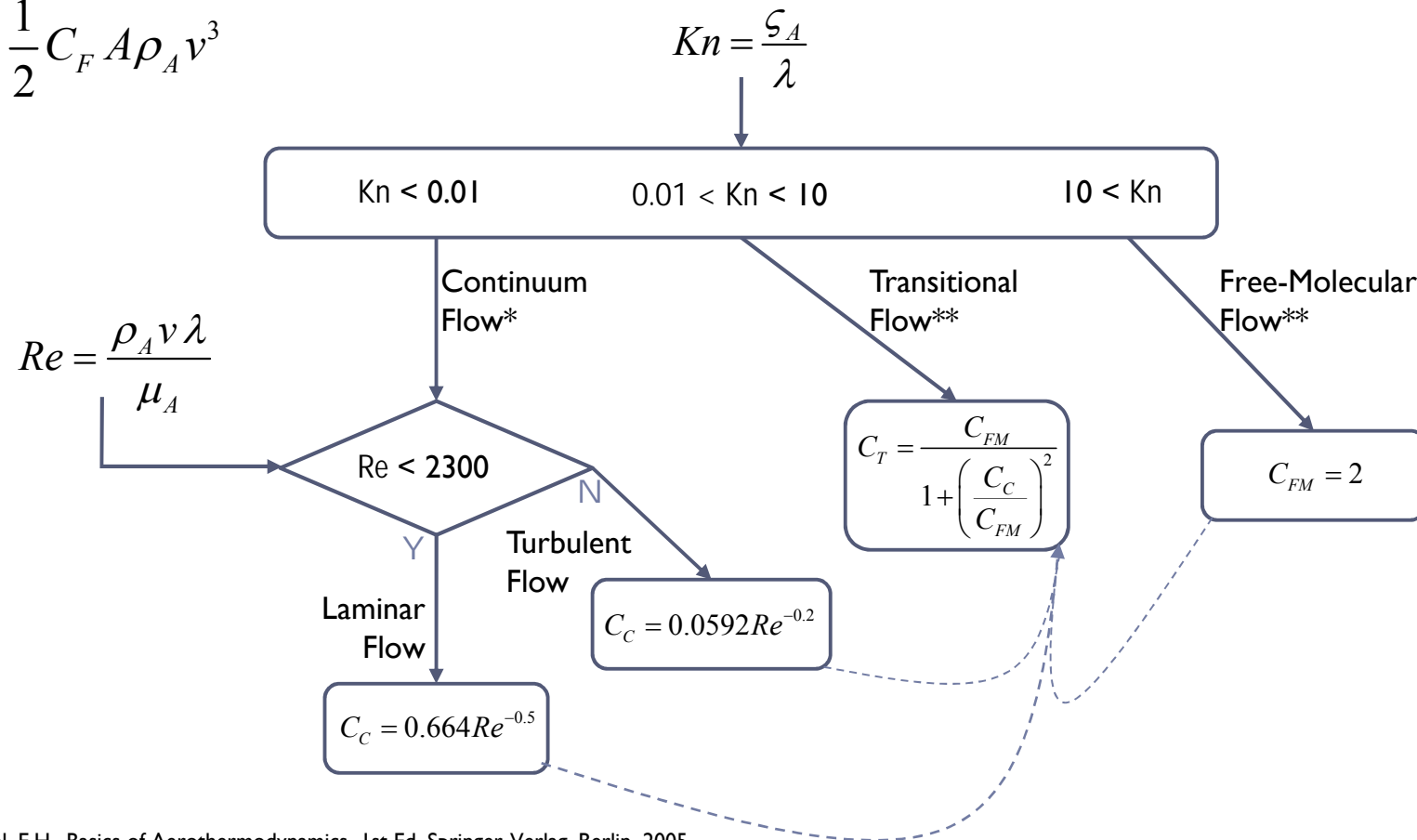
Sprite Features





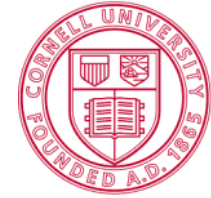
Skin Friction Coefficient

$$Q_c = \frac{1}{2} C_F A \rho_A v^3$$



* Hirshcel, E.H., Basics of Aerothermodynamics, 1st Ed, Springer-Verlag, Berlin, 2005.

**Koppenwallner, G., Fritsche, Band Bolts of Disintegrating Spacecraft during Uncontrolled re-entry", Proceedings of the Third European Conference on Space Debris, 19 - 21 March 2001, .. and T. Lips, "Survivability and Ground Risk Potential of Screws Darmstadt, Germany.



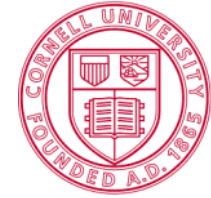
Sprite Power System

Parameter	Value	Units
Solar Cells	6	
Voltage	48	V
Current	50	μA
Capacitor	20	μF
Maximum Voltage	30	V
Minimum Voltage	18	V
Charge Time	1.0	s
Released Energy	6.4	mJ
Pulse Length	18	ms
DC-DC Efficiency	80	%
Filtered DC Output		
Supply Voltage	3.3	V
Supply Current	86	mA
Power Consumption		
Microcontroller Power	-0.6	mA
Transmitter Power	-29	mA
Margin	56	mA

Sprite Communications Subsystem

Cornell University

Space Systems Design Studio



Parameter	Value	Units
Transmitter		
Power	10	mW
Antenna Gain	0	dB
Frequency	902	MHz
Bandwidth	1	MHz
Pulse Length	16	ms
Chip Rate	10	kHz
Orbit		
Altitude	500	km
Overhead Arc-Length	+/- 20	deg
Free Space Loss	-146	dB
Atmospheric Attenuation	-2	dB
Receiver		
Receiver Gain	18	dB
Noise Temperature	300	K
Noise Power	-142	dBW
Margin	5	dB
Received Power	-154	dBW
Carrier to Noise Ratio	0.07	
Matched Filter Gain	22	dB
Signal to Noise Ratio	11	

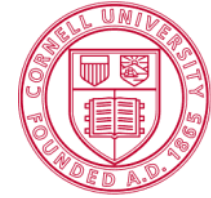
Concept

Design

Simulation

Conclusions

↳ Backup Slides

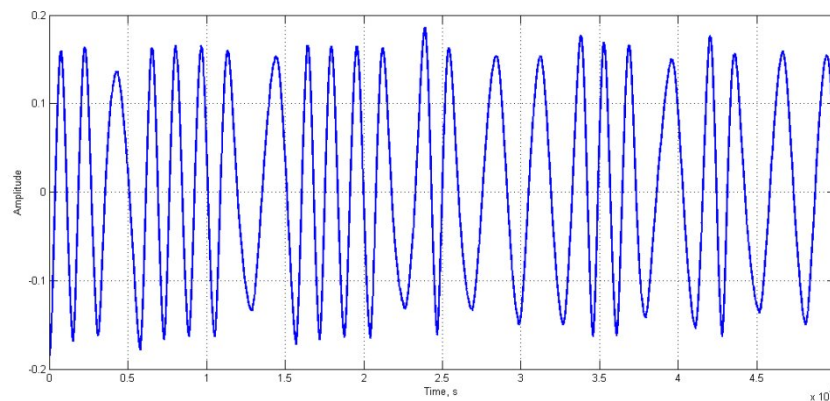


Communications

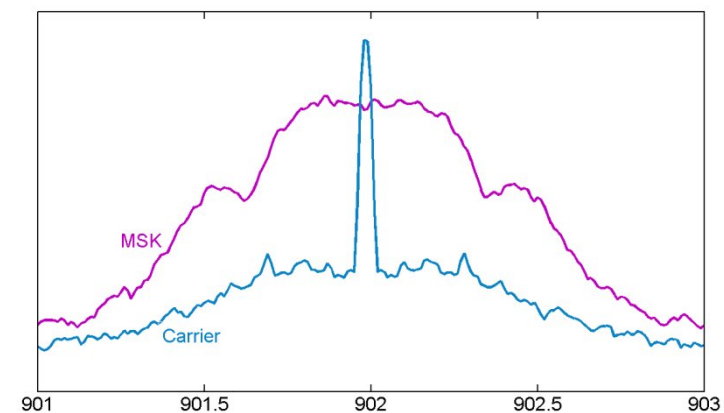
“Minimum Shift Keying” (MSK) Signal Modulation

MSK has the advantage of efficiently spreading the signal's energy over a range of frequencies.

MSK Modulated Signal



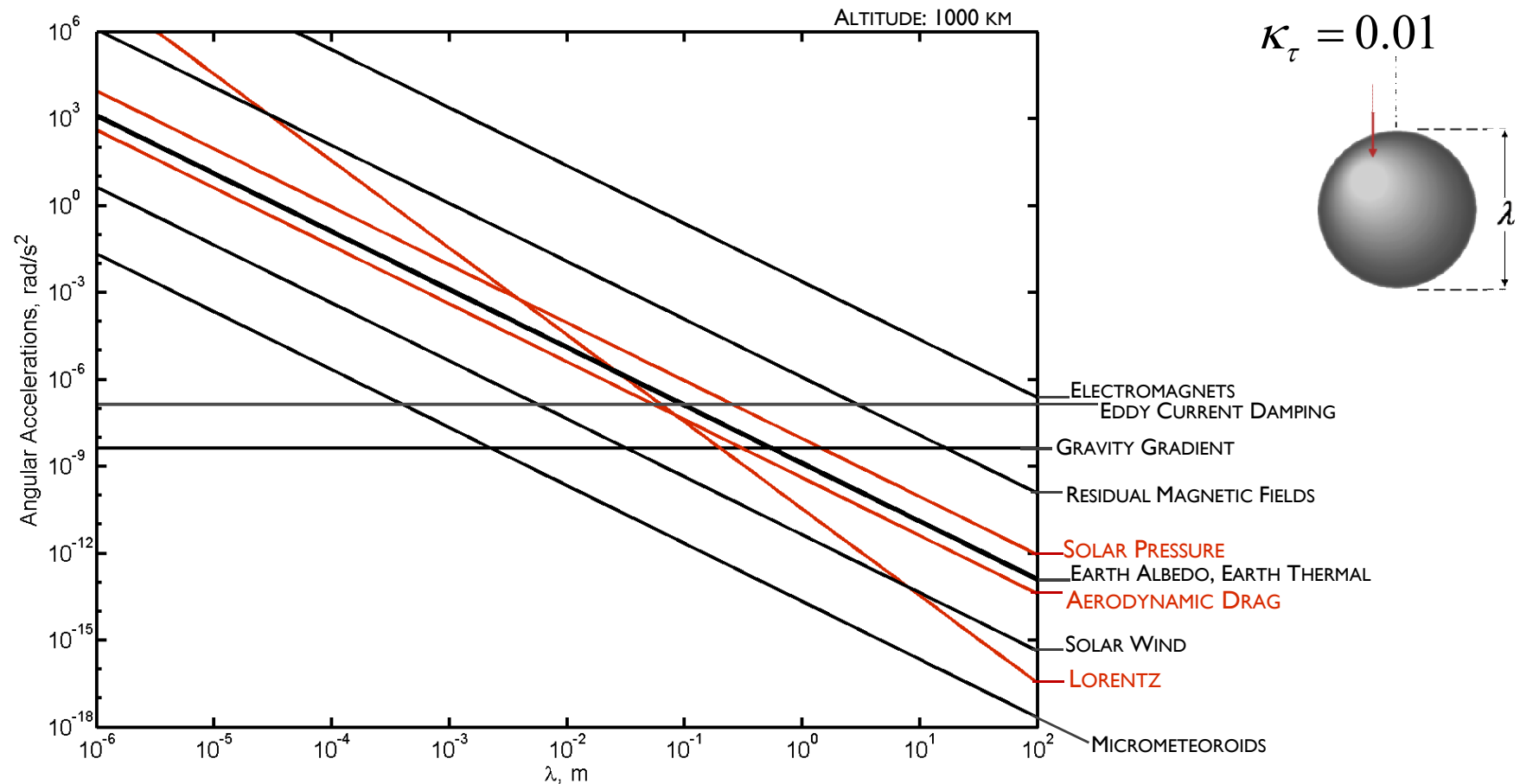
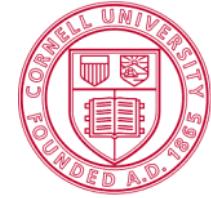
Power Spectral Density



Angular Accelerations and Length Scale

Cornell University

Space Systems Design Studio



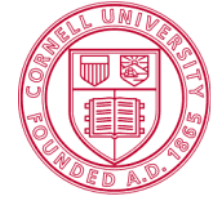
Concept

Design

Simulation

Conclusions

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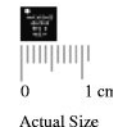
Small Sensors

Miniaturization
seems to be
today's design
paradigm.

“Mass is King”



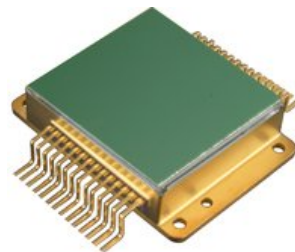
CMOS Imager



InvenSense 2-Axis Gyro



Kionix 3-Axis Accelerometer



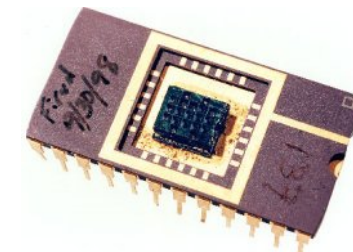
ULIS Microbolometer



Rockland
Micromagnetometer



AeroAstro Medium Sun Sensor



DARPA Digital Micropropulsion